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A MATLAB-based graphical user interface program for computing functionals of the geopotential up to ultra-high degrees and orders: efficient computation at irregular surfaces[☆]

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Abstract

1 Fast spherical harmonic synthesis (SHS) at multiple points based on the lumped
2 coefficients approach is very well-established technique. However, this method
3 cannot be applied to SHS at irregular surfaces, as the points must be regularly
4 spaced and refer to a regular surface such as the sphere or the ellipsoid of rev-
5 olution. In this paper we present a MATLAB[®]-based graphical user interface
6 software for ultra-high degree SHS on grids at irregular surfaces, like the Earth
7 surface. This software employs the highly efficient lumped coefficient approach
8 for SHS at regular surfaces and the Taylor series expansions to continue the func-
9 tionals to the irregular surfaces, e.g. the Earth surface. The generalized idea of
10 continuing functionals using the Taylor series was presented by Hirt (Journal of
11 Geodesy 86:729–744, 2012). We took the advantage of the software GrafLab

[☆]Code available from http://www.svf.stuba.sk/en/departments/departments-of-theoretical-geodesy/science-and-research/downloads.html?page_id=4996

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12 (Bucha and Janák in Computers & Geosciences 56:186–196, 2013), which em-
13 ploys the lumped coefficients approach, and developed a new software isGrafLab
14 (Irregular Surface GRAvity Field LABoratory). Compared to the commonly used
15 “two loops” approach, the factor of increased computational speed can reach a
16 value of several hundreds. isGrafLab allows accurate evaluation of 38 functionals
17 of the geopotential on grids at irregular surfaces. High orders of the Taylor series
18 can be used for the continuation. The new software offers all the other options
19 available in GrafLab, such as the employment of three different approaches to
20 compute the fully normalized associated Legendre functions, the graphical user
21 interface or the possibility to depict data on a map.

22 *Keywords:* Geopotential model, Gravity field functional, Spherical harmonic
23 synthesis, Lumped coefficients, Irregular surface, Taylor series

24 **1. Introduction**

25 In geosciences, the spherical harmonics (e.g. Freeden and Schreiner, 2009;
26 Hobson, 1931; Hoffmann-Wellenhof and Moritz, 2005), which form orthonormal
27 basis functions with a global support on the sphere, are frequently used to repre-
28 sent scalar fields such as the Earth gravitational (e.g. Pavlis et al., 2012) or mag-
29 netic (e.g. Finlay et al., 2010) field. By means of the spherical harmonics, these
30 fields can be transformed from spatial domain into frequency domain (spherical
31 harmonic analysis (SHA), e.g. Colombo (1981); Sneeuw (1994)), in which they
32 are characterized by a set of spherical harmonic coefficients. These coefficients
33 can be transformed back into the spatial domain (spherical harmonic synthesis
34 (SHS), e.g. Barthelmes (2013); Fantino and Casotto (2009)), and in the case of
35 the gravitational field one gets the gravitational potential (in short geopotential).

36 After applying some operator to the spherical harmonics (e.g. the pseudodiffer-
37 ential operators (cf. Freeden and Schreiner, 2009)), commonly used functionals
38 of the geopotential can easily be obtained (e.g. gravitational tensor). When in-
39 troducing the normal gravity field (cf. Hoffmann-Wellenhof and Moritz, 2005),
40 functionals of the disturbing potential are easy to calculate as well (e.g. gravity
41 anomaly, deflections of the vertical, geoid undulation, etc.).

42 Nowadays, the gravity field of the Earth is very well-known up to a certain
43 level of frequencies, i.e. up to some degrees of spherical harmonics to which a
44 certain spatial resolution corresponds. Global geopotential models of the Earth
45 (GGMs) consist of spherical harmonic coefficients representing the Earth's grav-
46 itational field and of a few additional constants. Nowadays, GGMs are almost
47 routinely used in geodesy and other geosciences (climatology, geology, geomor-
48 phology, geophysics, oceanography, etc.). As high-resolution GGMs we would
49 like to mention EGM2008 (Pavlis et al., 2012) up to degree 2190, EIGEN-6C
50 (Förste et al., 2011) up to 1420 and EIGEN-6C2 (Förste et al., 2012) up to 1949.

51 Certainly the most difficult part of the SHS is the numerical evaluation of high
52 degree and order spherical harmonics, viz. the fully normalized associated Legen-
53 dre functions of the first kind (fnALFs) (see e.g. Hoffmann-Wellenhof and Moritz,
54 2005). Since fnALFs cover a huge range of magnitude, underflow and overflow
55 problems occur in their evaluation (Holmes and Featherstone, 2002). A lot of at-
56 tention has been put on accurate and fast computation of these functions in the
57 last decades. These efforts resulted in the development of several approaches to
58 compute fnALFs, differing mainly in numerical stability and efficiency (Balmino
59 et al., 2012; Fukushima, 2012a,b; Gruber et al., 2011; Holmes and Featherstone,
60 2002; Cheong et al., 2012; Jekeli et al., 2007; Nesvadba, 2009; Smith et al., 1981;

61 Šprlák, 2011; Wenzel, 1998).

62 In order to resolve the other problem of SHS, i.e. the numerical efficiency of
63 computation, one frequently uses the well-known lumped coefficients approach
64 (LCA), see Section 2.1 and the references therein. This method allows perform-
65 ing very fast SHS on the grids that are regularly spaced and refer to a regular
66 surface (the sphere or the ellipsoid of revolution). However, this approach cannot
67 be applied if one of these conditions is not fulfilled. Hirt (2012) recently presented
68 an effective method, in which one of these conditions, namely the regularity of the
69 surface, can be omitted. He generalized the approach published already by Rapp
70 (1997), who utilized the first-order Taylor series to continue the height anomaly
71 from the ellipsoid to the Earth surface. In other words, certain functional and its
72 derivatives of chosen order can be computed on the sphere or the ellipsoid em-
73 ploying the LCA, and subsequently continued to an irregular surface, e.g. the
74 Earth surface. This approach is very simple in principle, but it can also be highly
75 efficient and accurate, as it will be pointed out. Basically an identical approach
76 was employed in Balmino et al. (2012), Section 4.3 and Appendix B.

77 Various computer programs for computing functionals of the geopotential are
78 available, e.g. Adams and Swarztrauber (1997); Barthelmes (2003); Holmes and
79 Pavlis (2008); Janák and Šprlák (2006); Sanso and Sona (2001); Smith (1998);
80 Tscherning et al. (1983); Wieczorek (2012); Wittwer et al. (2008) or recently pub-
81 lished Bucha and Janák (2013). A software to compute elements of the disturbing
82 tensor on grids at irregular surfaces was presented by Eshagh and Abdollahzadeh
83 (2012). However, to the knowledge of the authors, there is no compact and pub-
84 licly available software with implemented accelerated routines that would allow
85 fast and accurate SHS on grids at irregular surfaces such as the Earth surface.

86 This obstacle one usually circumvents by the point-wise approach with two loops,
 87 which can be, however, slower by a factor of several hundreds. This led us to
 88 modify the software GrafLab 1.1.2 (Bucha and Janák, 2013) to provide fast and
 89 accurate ultra-high degree SHS of functionals at irregular surfaces. The new soft-
 90 ware is called isGrafLab (Irregular Surface GRAvity Field LABoratory). It is
 91 written in MATLAB[®] and equipped with an easy-to-use graphical user interface,
 92 therefore no modifications on the source code level are necessary, as it is in the
 93 vast majority of the software of this kind.

94 The paper is organized as follows. Section 2 provides a brief recapitulation of
 95 the lumped coefficient approach and description of the Taylor expansions in the
 96 context of SHS at irregular surfaces. Section 3 describes the basic principles of
 97 manipulation with the software, and assessment of the accuracy and efficiency of
 98 the developed program. Summary and conclusions are presented in Section 4.

99 **2. Spherical harmonic synthesis at irregular surfaces**

100 *2.1. Lumped coefficients approach*

101 Disturbing potential (see e.g. Hoffmann-Wellenhof and Moritz, 2005) expanded
 102 into the truncated series of the spherical harmonics and written in the form that
 103 employs the lumped coefficients (e.g. Colombo, 1981; Fantino and Casotto, 2009)
 104 reads

$$T(r, \theta, \lambda) = \frac{GM}{r} \sum_{m=0}^M (A_m(\theta) \cos m\lambda + B_m(\theta) \sin m\lambda), \quad (1)$$

105 where

$$A_m(\theta) = \sum_{n=m}^M \left(\frac{R}{r}\right)^n \bar{C}_{nm} \bar{P}_{nm}(\cos \theta), \quad (2a)$$

$$B_m(\theta) = \sum_{n=m}^M \left(\frac{R}{r}\right)^n \bar{S}_{nm} \bar{P}_{nm}(\cos \theta), \quad (2b)$$

and r, θ, λ denotes the spherical radius, the spherical co-latitude and the longitude, respectively, GM is the geocentric gravitational constant, R is the radius of the reference sphere, $\{\bar{C}_{nm}, \bar{S}_{nm}\}$ is the set of fully normalized spherical harmonic coefficients of degree n and order m (we assume that the spherical harmonic coefficients of the normal field have been subtracted, see e.g. Barthelmes (2013), Eqs. (109) and (111)), $\bar{P}_{nm}(\cos \theta)$ are the fully normalized associated Legendre functions of the first kind (fnALFs) of degree n and order m , M denotes the maximum degree of spherical harmonic expansion and, finally, $A_m(\theta)$ and $B_m(\theta)$ stand for the lumped coefficients of the order m .

The efficiency of the LCA is in the fact that these coefficients (see Eqs. (2a) and (2b)) are constant along parallels in the grid, i.e. independent of the longitude. Therefore if one uses the standard co-latitude–longitude grid at the sphere or the ellipsoid of revolution, they need to be computed only once. Another trick is that both summations, i.e. over m and subsequently over n , can be summed by matrix multiplications, which significantly reduces the computation time. For the full vector-matrix notation of Eq. (1) see e.g. Bucha and Janák (2013).

An arbitrary functional of the geopotential can be written in the form of the lumped coefficients, and thus significantly speeds up (in some cases by a factor of several hundreds) SHS at regular surfaces.

2.2. Gradient approach in spherical coordinates

From Section 2.1 it is clear that the LCA cannot be applied to SHS on grids at the Earth surface, as the topography is varying differently in each meridian, and thus the coefficients $A_m(\theta)$ and $B_m(\theta)$ are not constant. An accurate, effective and very simple approach how to circumvent this issue has been proposed and verified by Hirt (2012), see also Balmino et al. (2012). They used Taylor expansions

132 (gradient approach) to continue functionals from regular surface, at which the
 133 LCA is employed, to irregular surface.

134 Let $P(r, \theta, \lambda) \in \Sigma$ be points of the standard co-latitude–longitude grid ly-
 135 ing at the Earth surface Σ (or at an arbitrary irregular surface in general), and
 136 $P_0(r_0, \theta, \lambda) \in S$ be points of the same grid referring to the sphere S with the
 137 radius r_0 . To continue disturbing potential from the points $P_0(r_0, \theta, \lambda) \in S$ to
 138 $P(r, \theta, \lambda) \in \Sigma$, the truncated Taylor series can be written in the form

$$T(r, \theta, \lambda) = T_0(r_0, \theta, \lambda) + \sum_{k=1}^K \frac{1}{k!} \frac{\partial^k T}{\partial r^k} \Big|_{P_0(r_0, \theta, \lambda) \in S} (r - r_0)^k, \quad (3)$$

139 where K is the maximum order of the Taylor series. The k^{th} radial derivative of
 140 the disturbing potential can be obtained by the simple formula

$$\frac{\partial^k T(r, \theta, \lambda)}{\partial r^k} = (-1)^k \frac{GM}{r^{k+1}} \sum_{m=0}^M \left(A_m^{(k)} \cos m\lambda + B_m^{(k)} \sin k\lambda \right), \quad (4)$$

141 in which

$$A_m^{(k)} = \sum_{n=m}^M \left(\frac{R}{r} \right)^n \left(\prod_{i=1}^k (n+i) \right) \bar{C}_{nm} \bar{P}_{nm}(\cos \theta), \quad (5a)$$

$$B_m^{(k)} = \sum_{n=m}^M \left(\frac{R}{r} \right)^n \left(\prod_{i=1}^k (n+i) \right) \bar{S}_{nm} \bar{P}_{nm}(\cos \theta). \quad (5b)$$

143 These equations employ the highly efficient LCA. After easy calculations,
 144 commonly used functionals of the geopotential can be continued to irregular sur-
 145 faces using this scheme. Similar formulae for the gravity disturbance, the height
 146 anomaly and the deflections of the vertical can be found in Hirt (2012).

147 It is obvious that increasing the maximum order of the Taylor series will lead
 148 to better approximation of the functional. Until now, no approximation has been
 149 done, except for the truncation of the Taylor series at some finite order value K .
 150 However, as it will be pointed out in Section 3.2, in the “gradient approach in

151 spherical coordinates” sufficiently high order K ensures negligible approximation
 152 errors.

153 2.3. Gradient approach in ellipsoidal coordinates

154 An approximation is necessary, if one uses the ellipsoidal coordinates (φ, λ, h) ,
 155 i.e. the ellipsoidal latitude, the longitude and the ellipsoidal height, respectively,
 156 as input data instead of the spherical coordinates (r, θ, λ) . The ellipsoidal coordi-
 157 nates of points $P(\varphi, \lambda, h) \in \Sigma$ at the standard latitude–longitude grid can be easily
 158 transformed into the spherical coordinates $P(r^E, \theta^E, \lambda) \in \Sigma$ by the well-known for-
 159 mulae (e.g. Hoffmann-Wellenhof and Moritz, 2005). However, due to the various
 160 values of h used in the original grid, the transformed co-latitudes θ^E vary along
 161 ellipsoidal parallels. Thus, the LCA cannot be applied. To ensure that θ^E will be
 162 constant along the parallels, some ellipsoid E at the ellipsoidal height h_0 above
 163 the reference ellipsoid can be used (h_0 can be e.g. average value computed from
 164 the minimum and the maximum ellipsoidal heights of the grid points), so that af-
 165 ter the transformation one gets $P_0(r_0^E, \theta_0^E, \lambda) \in E$. The spherical co-latitudes θ_0^E
 166 are now constant along each parallel and the LCA can be employed. Functionals
 167 are therefore continued from the ellipsoidal surface $E \ni P_0(r_0^E, \theta_0^E, \lambda)$ to the Earth
 168 surface $\Sigma \ni P(r^E, \theta^E, \lambda)$.

169 The truncated Taylor expansion of disturbing potential reads

$$T(r^E, \theta^E, \lambda) \approx T_0(r_0^E, \theta_0^E, \lambda) + \sum_{k=1}^K \frac{1}{k!} \frac{\partial^k T}{\partial r^k} \Big|_{P_0(r_0^E, \theta_0^E, \lambda) \in E} (h - h_0)^k. \quad (6)$$

170 The approximate sign in Eq. (6) is used because of the gradients computed in
 171 the radial direction, which differ from the gradients in the direction of $h - h_0$.
 172 This is, however, acceptable, as it was shown in Hirt (2012) and will be shown
 173 in Section 3.2 as well. Unlike the “gradient approach in spherical coordinates”,

174 this also causes that beyond certain value of K , incessantly increasing the order
 175 of the Taylor series will not have the effect of permanently “better” and “better”
 176 approximation of the true value. In other words, the Taylor series converges, but
 177 to a value of the functional at slightly different point, and not to the true value.

178 2.4. Remark on the gradient approach of some functionals

179 Let us mention a short remark on the gradient approach of some function-
 180 als, e.g. the height anomaly, the components of deflection of the vertical, etc.
 181 These quantities are composite functions, because of the normal gravity γ (e.g.
 182 Hoffmann-Wellenhof and Moritz, 2005), which is also dependent on r . Thus, the
 183 truncated Taylor series for, let’s say, the height anomaly reads (“approach in the
 184 spherical coordinates”)

$$\zeta(r, \theta, \lambda) = \zeta_0(r_0, \theta, \lambda) + \sum_{k=1}^K \frac{1}{k!} \left. \frac{\partial^k T/\gamma}{\partial r^k} \right|_{P_0(r_0, \theta, \lambda) \in S} (r - r_0)^k. \quad (7)$$

185 When $k > 2$, the partial derivatives of the term T/γ significantly complicate
 186 numerical computation of Eq. (7), as both the disturbing potential T and the nor-
 187 mal gravity γ are dependent on r . A very simple way how to avoid this problem,
 188 which was used in isGrafLab, is that the functional and its derivatives can be
 189 computed on the grid without the term $\gamma(r_0, \theta)$, and subsequently continued to the
 190 Earth surface. After this, the intermediate result is divided by $\gamma(r, \theta)$ and the fi-
 191 nal result is obtained. This method is faster than evaluating the high-order partial
 192 derivatives of the term T/γ . The same holds for the functionals, in which the terms
 193 like $1/\sin \theta$ occur or contain the centrifugal potential or acceleration (although the
 194 reasons are slightly different here). On the other hand, the price to be paid is the
 195 cost of additional memory, since full matrix of $\gamma(r, \theta)$ (or $1/\sin \theta^E$ in the “gradient

196 approach in ellipsoidal coordinates”) has to be stored. The same approach was
197 used for the height anomaly in Balmino et al. (2012), Appendix B.

198 **3. Software presentations**

199 *3.1. Description of the software*

200 The aim of this section is to present the new developed software isGrafLab
201 (Irregular Surface GRAvity Field LABoratory), which is modified version of Graf-
202 Lab 1.1.2 (Bucha and Janák, 2013), and to assess the effectiveness of the gradient
203 approach in spherical and ellipsoidal coordinates. isGrafLab, as well as GrafLab,
204 is written in MATLAB® and equipped with a graphical user interface (GUI), see
205 Fig. 1.

206 Fig. 1 should be positioned here

207 The GUI is visually divided into three panels, from which the first and the third
208 panel were changed only slightly. In order to stay brief, for the details about the
209 manipulation with these two panels we would like to advise the reader to Bucha
210 and Janák (2013), Section 3.1.

211 (i) *Geopotential model and reference system selection*: The only change in
212 this panel is that isGrafLab is not capable to read a variance-covariance matrix of
213 spherical harmonic coefficients as the input file. This option was left out, since
214 the main problem in the computation of commission error is not in the slowness
215 of the computation as much as in the high requirements on the RAM of computer.
216 Thus, the point-wise approach in GrafLab is certainly sufficient.

217 (ii) *Irregular surface selection*: At first, the *Type of the input coordinates* (el-
218 lipsoidal/spherical) must be specified. The grid must be defined by using the seven

219 self-explanatory arrays in the bottom left of this panel (*Lat.* denotes the latitude
 220 and *Lon.* is the longitude). The array *Height above the reference surface (m)* de-
 221 notes the constant height of the grid above the reference ellipsoid (GRS80 (Moritz,
 222 2000) or WGS84 (NIMA, 2000)) in the case of the ellipsoidal type of the coordi-
 223 nates or above the reference sphere with the radius R , defined by the GGM, in the
 224 case of the spherical coordinates. To this surface refer the points $P_0(r_0^E, \theta_0^E, \lambda) \in E$
 225 or $P_0(r_0, \theta, \lambda) \in S$ (see Secs. 2.3 and 2.2, respectively) that will be continued to the
 226 irregular surface. By a rule of thumb, to this array one can enter the average value
 227 computed from the minimum and the maximum height of the irregular surface
 228 above the reference surface (the ellipsoid GRS80/WGS84 or the sphere with the
 229 radius R). The entries must be either floating point numbers with decimal points
 230 or integer values. Latitudes must be entered within the $\langle -90^\circ, 90^\circ \rangle$ interval and
 231 longitudes within the $\langle 0^\circ, 360^\circ \rangle$ or $\langle -180^\circ, 180^\circ \rangle$ interval.

232 An integer value of the maximum order of the Taylor series ($K \geq 0$) (see
 233 Eq. (3)) must be entered into the array *Order of Taylor series*:. If $K = 0$, the
 234 functionals are not continued from the regular surface.

235 Heights of the grid points at the irregular surface must be loaded using the
 236 *Browse...* button. In case of the ellipsoidal input coordinates this file must contain
 237 the ellipsoidal heights. If the spherical coordinates have been chosen, the spherical
 238 radii of the grid points must be given. In both cases, the input file may have two
 239 forms:

240 a) a matrix

$$\begin{bmatrix} (\theta_n, \lambda_1) & \cdots & (\theta_n, \lambda_m) \\ \vdots & \ddots & \vdots \\ (\theta_1, \lambda_1) & \cdots & (\theta_1, \lambda_m) \end{bmatrix},$$

241 b) a column vector (note the transposition)

$$\begin{bmatrix} (\theta_1, \lambda_1) & \dots & (\theta_n, \lambda_1) & \dots & (\theta_1, \lambda_m) & \dots & (\theta_n, \lambda_m) \end{bmatrix}^T,$$

242 where (θ_1, λ_1) are the coordinates of the southernmost and the westernmost grid
243 point (arrays *Lat. min* (°) and *Lon. min* (°)). The northernmost and the easternmost
244 grid point is denoted by (θ_n, λ_m) (arrays *Lat. max* (°) and *Lon. max* (°)). The input
245 file may either be an ASCII file or a binary MAT-file. Points with undefined
246 ellipsoidal heights/spherical radii should be indicated by the value −9999 or NaN.
247 The output values are assigned correspondingly.

248 (iii) *Calculated parameters and output selection*: From this panel only the
249 check-box *Commission error* was removed. The rest of the panel was left un-
250 changed. For the convenience of the reader, let us herein list again the summary
251 of the functionals that can be computed in isGrafLab, see Table 1. We also note
252 that the formulae to compute each functional can be found in the pdf file *Defini-*
253 *tion_of_functionals_of_the_geopotential_used_in_GrafLab_software.pdf* available at ¹.

254 Table 1 should be positioned here

255 Here, a short note is necessary. The approaches for computing the functionals
256 *Geoid undulation* and *Height anomaly* are the same in isGrafLab as in GrafLab.
257 The reason is that these functionals are referred to particular surfaces, and can
258 not be, by definition, shifted arbitrarily in radial or vertical direction. Moreover,
259 these functionals are computed iteratively as the particular surfaces are not known
260 prior to computation. However, for the convenience of the users we left these

¹http://www.svf.stuba.sk/en/departments/departments-of-theoretical-geodesy/science-and-research/downloads.html?page_id=4996

261 functionals also in isGrafLab.

262 isGrafLab enables to depict computed data on a map using the Mapping toolboxTM
263 of MATLAB[®]. For visualizing data using only the basic module of MATLAB[®],
264 see e.g. Bezděk and Sebera (2013).

265 When all the required input parameters and input files have been entered, after
266 clicking the *OK* button, the computation will start. On the left from this button,
267 there is a status line, which provides short explanations during the whole com-
268 putational process (*Loading GGM file...*, current value of the variable m in the
269 order-dependent loop, *Displaying data...*, etc.), so that a user can clearly see in
270 which part of the computation the isGrafLab is. After successful computation, the
271 status *Computation has been finished* will appear. If any of the input parameters
272 or input files have been entered in a wrong format, isGrafLab will open a warning
273 dialog or error dialog with description of the error.

274 The whole source code of the program called isGrafLab.m is written as a one
275 m-file. A simple flowchart of the program is provided in Fig. 2.

276 Fig. 2 should be positioned here

277 3.2. Testing of the gradient approach

278 The assessment of numerical results obtained from isGrafLab is straightfor-
279 ward. In a test area we computed chosen functionals in isGrafLab with increasing
280 the maximum order of the Taylor series, and compared them to the functionals
281 obtained from GrafLab 1.1.2 using the point-wise approach. These values were
282 assumed to be the reference values at the same points. Differences between the
283 reference values and those obtained from the gradient approach indicate the level
284 of approximation of this method.

285 Fig. 3 should be positioned here

286 Let us start with the “gradient approach in ellipsoidal coordinates” at first. For
287 better comparison we will closely follow the test designed by Hirt (2012) with
288 only small modifications. As the test area we used the roughest region on the
289 globe, viz. the part of the Himalayas ($\varphi \in (26^\circ, 29^\circ)$, $\lambda \in (86^\circ, 88^\circ)$, 1’ spatial res-
290 olution) that includes the Mount Everest, see Fig. 3. The topography within this
291 area shows the largest variations on the Earth surface from ~ 0 km to ~ 8.8 km. In
292 order to obtain the grid of 180×120 points $P(\varphi, \lambda, h) \in \Sigma$ with ellipsoidal heights
293 of the topography within this area, we used SRTM (Shuttle Radar Topography
294 Mission) digital elevation model Version 4.1 (Jarvis et al., 2008) and the quasi-
295 geoid from EGM2008 (evaluated up to degree 2190 in GrafLab 1.1.2). Note that
296 the SRTM elevations are referred to EGM96 (Lemoine et al., 1998). However, to
297 achieve preferably similar results as in the test by Hirt, we used EGM2008 (see
298 Hirt, 2012, Section 2.4.1). The impact of the choice of the reference surface E
299 to the continued functionals is also described in Hirt (2012). Based on his in-
300 vestigations, herein we used as the regular surface E only the “best” option, i.e.
301 the ellipsoid at the height $h_0 = (\min(h) + \max(h))/2 \approx 4$ km above the reference
302 ellipsoid. At this grid represented by 21 600 points we computed δg_{sa} and T_{rr} for
303 $K = 0, \dots, 7$ using EGM2008 to degree 2190, and continued them to the Earth
304 surface Σ . Compared to Hirt (2012), in our test we extended the order of the Tay-
305 lor series from $K = 3$ to $K = 7$ (note that in isGrafLab the value of K can be even
306 larger). The statistical characteristics of this test are shown in Table 2. Fig. 4 gives
307 a graphical representation of the differences for δg_{sa} . The computations were ex-
308 ecuted on a PC with Intel® Core™ i5-3330 CPU and 4 GB RAM under the 64 bit
309 Windows® 7 OS.

310 Fig. 4 should be positioned here

311 Table 2 should be positioned here

312 Table 2 shows very similar characteristics for δg_{sa} and $K \leq 3$ to those by Hirt
313 (2012), Table 4. It is also obvious that by further increasing the maximum order of
314 the Taylor series beyond $K = 3$, the better approximation of the reference values
315 can be achieved. However, as it was already pointed out in Section 2.3, this con-
316 vergence has to “stop” at some value of K . This is caused by the approximation
317 of the gradients in the direction of $h - h_0$ by the gradients in the radial direction,
318 see Section 2.3. For $M = 2190$ and the chosen region, the convergence “stops” at
319 $K \approx 5$. However, positive effect of the increased K beyond 3 is evident. For δg_{sa}
320 the RMS fell from 0.23 mGal ($K = 3$) to 0.03 mGal ($K = 5$), and from 0.73 E
321 ($K = 3$) to 0.08 E ($K = 5$) for T_{rr} , which can be considered as significant im-
322 provement. These errors are far bellow the commission errors of EGM2008 (Hirt,
323 2012; Pavlis et al., 2012), therefore this approach can be considered as accurate
324 even in extreme conditions. It should be emphasized that for smaller values of M
325 and/or in smoother areas the results are even more optimistic.

326 Now let us focus on the “gradient approach in spherical coordinates”, in which
327 the convergence never “stops” essentially, since no approximation is used. The
328 ellipsoidal coordinates of the same points $P(\varphi, \lambda, h) \in \Sigma$ within the same area as
329 in the test above were transformed into the spherical coordinates $P(r^E, \theta^E, \lambda) \in$
330 Σ . To ensure that the transformed θ^E will be constant along the parallels and
331 subsequently the LCA will be usable, we let $\theta^E = 90^\circ - \varphi$ and obtain the regular
332 grid points $P(r^E, \theta^E, \lambda) \in \Sigma = P(r^E, 90^\circ - \varphi, \lambda) \in \Sigma$. This approximation is
333 admissible, since we are testing only the method herein. Next, we again computed
334 at the points $P(r_0^E, \theta^E, \lambda) \in S$ the functionals δg_{sa} , T_{rr} and their derivatives using

EGM2008 to degree 2190, and continued them from the regular surface S to the surface Σ . The radius r_0^E of the sphere S (see Eq. (3)) was obtained by $r_0^E = (\min(r^E) + \max(r^E))/2 \approx R = 6\,378\,136.3$ m. The order of the Taylor series was set to $K = 0, \dots, 20$. The results of this test are reported in Table 3.

Table 3 should be positioned here

It is clear that the Taylor series converges now to the reference values, although the convergence “slows down” at $K \approx 15$. In the case that $K = 15$ the RMS is 1.5×10^{-11} mGal for δg_{sa} and 4.7×10^{-11} E for T_{rr} . For smaller values of M the Taylor series converges even faster. Based on these investigations it can be said that for sufficiently high order of the Taylor series the approximation errors of the “gradient approach in spherical coordinates” are negligible.

3.3. Testing of the time efficiency

In this section we assess the time efficiency of the presented software isGrafLab. We compared computation time of the disturbing tensor in the local north-oriented reference frame (LNOF) using isGrafLab and the software presented by Eshagh and Abdollahzadeh (2012). This software is also written in MATLAB[®] and computes elements of the disturbing tensor at irregular surfaces. In addition, these elements can be evaluated in three different reference frames.

The test was executed on an ordinary PC with Intel[®] Core[™] i5-3330 CPU and 4 GB RAM under the 64 bit Windows[®] 7 OS. We used EGM2008 to degree 2190 and a global grid at $5' \times 5'$ spatial resolution which corresponds to 9 337 681 points. In isGrafLab we used two sufficiently high orders of the Taylor series (see Table 2), $K = 3$ and $K = 5$. We also applied two of the three available approaches to compute fnALFs in the software (for more details, see Bucha and

359 Janák, 2013, Section 2.1), namely the standard forward column method (Holmes
360 and Featherstone, 2002) and the extended-range arithmetic (Fukushima, 2012a).
361 The results of the test are summarized in Table 4.

362 Table 4 should be positioned here

363 Table 4 reveals that isGrafLab is capable to synthesize the disturbing ten-
364 sor about 20–10 times faster than the software presented by Eshagh and Abdol-
365 lahzadeh (2012). On the other hand, their approach computes the disturbing tensor
366 without any approximations, unlike the “gradient approach in *ellipsoidal* coordi-
367 nates”.

368 4. Summary

369 The novel developed software isGrafLab (Irregular Surface GRAvity Field
370 LABoratory) has been presented in this paper. isGrafLab is capable to perform
371 fast ultra-high degree SHS at regular spaced grids referring to irregular surfaces,
372 e.g. the Earth surface. The approach implemented in the software was published
373 by Hirt (2012) and also by Balmino et al. (2012). At first, functionals and their
374 derivatives are computed at a regular surface (the sphere or the ellipsoid of revolu-
375 tion) by means of the efficient LCA. Subsequently, the functionals are continued
376 to the irregular surface by applying the Taylor expansion. The order of the Tay-
377 lor series can be entered by the user. Compared to the “point-wise approach”
378 with two loops that is usually used in this situation, isGrafLab allows accurate
379 and time-saving SHS at dense grids at the Earth surface. The factor of increased
380 computational speed can reach a value of several hundreds. Note that depending
381 on the order of the Taylor series and the number of grid points, in general this

382 approach involves higher requirements on the RAM of computer compared to the
383 LCA alone. isGrafLab is written in MATLAB[®] and equipped with an easy-to-use
384 GUI. isGrafLab allows SHS of 38 functionals of the geopotential employing three
385 different approaches to compute fnALFs with various numerical stability and ef-
386 ficiency (see Bucha and Janák, 2013, Section 2.1).

387 The numerical investigations have been divided into two separate parts, namely
388 the gradient approach in ellipsoidal and spherical coordinates. It was shown that
389 in the former approach the RMS of the approximation errors can be decreased
390 (compared to Hirt (2012)) by further increasing the maximum order of the Taylor
391 series K . For $K = 5$, the maximum degree $M = 2190$ and the roughest topography
392 on the Earth, viz. the Himalayas, the RMS of the approximation errors reached
393 0.03 mGal for δg_{sa} and 0.08 E for T_{rr} . As for the later approach, it was demon-
394 strated that sufficiently large order of the Taylor series ensures agreement with the
395 reference values with negligible approximation errors.

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403 **References**

- 404 Adams, J.C., Swarztrauber, P.N., 1997. SPHEREPACK 2.0: A Model Develop-
405 ment Facility. NCAR Technical Note NCAR/TN-436-STR, 58 pp.
- 406 Balmino, G., Vales, N., Bonvalot, S., Briais, A., 2012. Spherical harmonic mod-
407 elling to ultra-high degree of Bouguer and isostatic anomalies. *Journal of*
408 *Geodesy* 86, 499–520. doi: 10.1007/s00190-011-0533-4.
- 409 Barthelmes, F., 2003. Calculation Service under ICGEM (International Centre for
410 Global Earth Models). <http://icgem.gfz-potsdam.de/ICGEM/>, (accessed
411 7 February, 2013).
- 412 Barthelmes, F., 2013. Definition of functionals of the geopotential and their calcu-
413 lation from spherical harmonic models: Theory and formulas used by the cal-
414 culation service of the International Centre for Global Earth Models (ICGEM),
415 <http://icgem.gfz-potsdam.de>. Scientific Technical Report STR09/02.
416 GFZ German Research Centre for Geosciences. Potsdam, Germany, 32 pp.
- 417 Bezdě́k, A., Sebera, J., 2013. Matlab script for 3D visualizing geodata
418 on a rotating globe. *Computers and Geosciences* 56, 127–130. doi:
419 10.1016/j.cageo.2013.03.007.
- 420 Bucha, B., Janák, J., 2013. A MATLAB-based graphical user interface
421 program for computing functionals of the geopotential up to ultra-high
422 degrees and orders. *Computers and Geosciences* 56, 186–196. doi:
423 10.1016/j.cageo.2013.03.012.
- 424 Cheong, H.B., Park, J.R., Kang, H.G., 2012. Fourier-series representation and

- 425 projection of spherical harmonic functions. *Journal of Geodesy* 86, 975–990.
426 doi: 10.1007/s00190-012-0558-3.
- 427 Colombo, O.L., 1981. Numerical methods for harmonic analysis on the sphere.
428 Report No. 310. Department of Geodetic Science and Surveying, The Ohio
429 State University. Columbus, Ohio, 140 pp.
- 430 Eshagh, M., Abdollahzadeh, M., 2012. Software for generating gravity gradients
431 using a geopotential model based on an irregular semivectorization algorithm.
432 *Computers and Geosciences* 39, 152–160. doi: 10.1016/j.cageo.2011.06.003.
- 433 Fantino, E., Casotto, S., 2009. Methods of harmonic synthesis for global geopo-
434 tential models and their first-, second- and third-order gradients. *Journal of*
435 *Geodesy* 83, 595–619. doi: 10.1007/s00190-008-0275-0.
- 436 Finlay, C.C., Maus, S., Beggan, C.D., Bondar, T.N., Chambodut, A., Cher-
437 nova, T.A., Chulliat, A., Golovkov, V.P., Hamilton, B., Hamoudi, M., Holme,
438 R., Hulot, G., Kuang, W., Langlais, B., Lesur, V., Lowes, F.J., Lühr, H.,
439 Macmillan, S., Mande, M., McLean, S., Manoj, C., Menvielle, M., Michaelis,
440 I., Olsen, N., Rauberg, J., Rother, M., Sabaka, T.J., Tangborn, A., Tøffner-
441 Clausen, L., Thébaud, E., Thomson, A.W.P., Wardinski, I., Wei, Z., Zvereva,
442 T.I., 2010. International geomagnetic reference field: The eleventh genera-
443 tion. *Geophysical Journal International* 183, 1216–1230. doi: 10.1111/j.1365-
444 246X.2010.04804.x.
- 445 Förste, C., Bruinsma, S., Shako, R., Marty, J.C., Flechtner, F., Abrikosov, O.,
446 Dahle, C., Lemoine, J.M., Neumayer, H., Biancale, R., Barthelmes, F., König,
447 R., Balmino, G., 2011. EIGEN-6 A new combined global gravity field model

448 including GOCE data from the collaboration of GFZ-Potsdam and GRGS-
 449 Toulouse, in: EGU General Assembly, Vienna, Austria, 3-8 April.

450 Förste, C., Bruinsma, S.L., Flechtner, F., Marty, J.C., Lemoine, J.M., Dahle, C.,
 451 Abrikosov, O., Neumayer, H., Biancale, R., Barthelmes, F., Balmino, G., 2012.
 452 A new release of EIGEN-6C, in: AGU Fall Meeting, San Francisco, USA, 3-7
 453 December.

454 Freeden, W., Schreiner, M., 2009. Spherical Functions of Mathematical Geo-
 455 sciences: A Scalar, Vectorial, and Tensorial Setup. Springer-Verlag, Berlin
 456 Heidelberg, 602 pp.

457 Fukushima, T., 2012a. Numerical computation of spherical harmonics of arbitrary
 458 degree and order by extending exponent of floating point numbers. Journal of
 459 Geodesy 86, 271–285. doi: 10.1007/s00190-011-0519-2.

460 Fukushima, T., 2012b. Numerical computation of spherical harmonics of arbitrary
 461 degree and order by extending exponent of floating point numbers: II first-,
 462 second-, and third-order derivatives. Journal of Geodesy 86, 1019–1028. doi:
 463 10.1007/s00190-012-0561-8.

464 Gruber, C., Novák, P., Sebera, J., 2011. FFT-based high-performance spherical
 465 harmonic transformation. Studia Geophysica et Geodaetica 55, 489–500. doi:
 466 10.1007/s11200-011-0029-y.

467 Hirt, C., 2012. Efficient and accurate high-degree spherical harmonic synthesis
 468 of gravity field functionals at the Earth’s surface using the gradient approach.
 469 Journal of Geodesy 86, 729–744. doi: 10.1007/s00190-012-0550-y.

- 470 Hobson, E.W., 1931. The Theory of Spherical and Ellipsoidal Harmonics. Cam-
471 bridge University Press, Cambridge, Great Britain, 500 pp.
- 472 Hoffmann-Wellenhof, B., Moritz, H., 2005. Physical Geodesy. Springer, Wien,
473 New York, 403 pp.
- 474 Holmes, S.A., Featherstone, W.E., 2002. A unified approach to the Clenshaw
475 summation and the recursive computation of very high degree and order nor-
476 malised associated Legendre functions. *Journal of Geodesy* 76, 279–299. doi:
477 10.1007/s00190-002-0216-2.
- 478 Holmes, S.A., Pavlis, N.K., 2008. Spherical harmonic synthesis software Har-
479 monic_synth. [http://earth-info.nga.mil/GandG/wgs84/gravitymod/
480 new_egm/new_egm.html](http://earth-info.nga.mil/GandG/wgs84/gravitymod/new_egm/new_egm.html), (accessed 5 December, 2012).
- 481 Janák, J., Šprlák, M., 2006. New software for gravity field modelling using spher-
482 ical harmonics. *Geodetic and Cartographic Horizon* 52, 1–8 (in Slovak).
- 483 Jarvis, A., Reuter, H.I., Nelson, A., Guevara, E., 2008. Hole-filled SRTM for
484 the globe Version 4. Available from the CGIAR-CSI SRTM 90m Database
485 (<http://srtm.csi.cgiar.org>).
- 486 Jekeli, C., Lee, J.K., Kwon, J.H., 2007. On the computation and approximation of
487 ultra-high-degree spherical harmonic series. *Journal of Geodesy* 81, 603–615.
488 doi: 10.1007/s00190-006-0123-z.
- 489 Lemoine, F.G., Kenyon, S.C., Factor, J.K., Trimmer, R.G., Pavlis, N.K.,
490 Chinn, D.S., Cox, C.M., Klosko, S.M., Luthcke, S.B., Torrence, M.H.,
491 Wang, Y.M., Williamson, R.G., Pavlis, E.C., Rapp, R.H., Olson, T.R., 1998.

492 The Development of the Joint NASA GSFC and the National Imagery and
 493 Mapping Agency (NIMA) Geopotential Model EGM96. Technical Report
 494 NASA/TP1998206861. National Aeronautics and Space Administration. USA,
 495 582 pp.

496 Moritz, H., 2000. Geodetic reference system 1980. *Journal of Geodesy* 74, 128–
 497 162. doi: 10.1007/s001900050278.

498 National Imagery and Mapping Agency (NIMA), 2000. *World Geodetic System*
 499 *1984: Its Definition and Relationships with Local Geodetic Systems*. Technical
 500 Report No. NIMA TR8350.2. National Imagery and Mapping Agency. USA,
 501 175 pp.

502 Nesvadba, O., 2009. Numerical problems in evaluating high degree and order
 503 associated Legendre functions, in: *General Assembly of the European Geo-*
 504 *sciences Union*, Vienna, Austria, 19-24 April.

505 Pavlis, N.K., Holmes, S.A., Kenyon, S.C., Factor, J.K., 2012. The development
 506 and evaluation of the Earth Gravitational Model 2008 (EGM2008). *Journal of*
 507 *Geophysical Research* 117(B04406), 1–38. doi: 10.1029/2011JB008916.

508 Rapp, R.H., 1997. Use of potential coefficient models for geoid undula-
 509 tion determinations using a spherical harmonic representation of the height
 510 anomaly/geoid undulation difference. *Journal of Geodesy* 71, 282–289.

511 Sanso, F., Sona, G., 2001. ELGRAM: an Ellipsoidal Gravity Model Manipulator.
 512 *Bollettino di Geodesia e Scienze Affini* 60, 3, 215–226.

513 Smith, D.A., 1998. There is no such thing as “the” EGM96 geoid: subtle points
 514 on the use of a global geopotential model. *IGeS Bulletin* 8, 17–28.

- 515 Smith, J.M., Olver, F.W.J., Lozier, D.W., 1981. Extended-range arithmetic and
516 normalized Legendre polynomials. *ACM Trans Math Softw* 7, 93–105.
- 517 Sneeuw, N., 1994. Global spherical harmonic analysis by least-squares and nu-
518 merical quadrature methods in historical perspective. *Geophysical Journal In-*
519 *ternational* 118, 707–716.
- 520 Tscherning, C.C., Rapp, R.H., Goad, C., 1983. A comparison of methods for com-
521 puting gravimetric quantities from high degree spherical harmonic expansions.
522 *Manuscripta Geodaetica* 8, 249–272.
- 523 Šprlák, M., 2011. On the numerical problems of spherical harmonics: numerical
524 and algebraic methods avoiding instabilities of associated Legendre’s functions.
525 *ZFV - Zeitschrift für Geodäsie, Geoinformation und Landmanagement* 136,
526 310–320.
- 527 Wenzel, G., 1998. Ultra-high degree geopotential models GPM98A, B and C to
528 degree 1800, in: Joint meeting of the International Gravity Commission and
529 International Geoid Commission, Trieste, Italy, 7-12 September.
- 530 Wieczorek, M., 2012. SHTOOLS – Tools for working with spherical harmonics.
531 <http://shtools.ipgp.fr/>, (accessed 28 November, 2013).
- 532 Wittwer, T., Klees, R., Seitz, K., Heck, B., 2008. Ultra-high degree spherical
533 harmonic analysis and synthesis using extended-range arithmetic. *Journal of*
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Table 1: Functionals of the geopotential available in isGrafLab. Explanation of the symbols in the table: V – gravitational potential, W – gravity potential, g – gravity, T – disturbing potential, δg – gravity disturbance, Δg – gravity anomaly, ξ – north-south component of deflection of the vertical, η – east-west component of deflection of the vertical, Θ – total deflection of the vertical, N – geoid undulation, ζ_{Ell} – generalized height anomaly, ζ – height anomaly; the subscript sa denotes the spherical approximation of the functional; (r, θ, λ) stands for the spherical coordinates; (x, y, z) denotes the coordinates in the local north-oriented reference frame; the subscripts $r, \theta, \lambda, x, y, z$ and their combinations stand for the derivatives of the functionals with respect to the particular coordinate; the number in the superscript denotes computational demand (computation time of the functional and memory usage during the computation) – (1) small, (2) medium, (3) high, (4) very high; (*) and (**) denote the functionals for which the value of n_{\min} cannot be larger than 2 and 0, respectively.

| Actual field | Disturbing field | Geometrical characteristics of the actual field |
|--|--|---|
| $V^{(1)}$ | $T^{(1)}$ | $\xi^{(2)}$ |
| $V_{rr}^{(3)}, V_{\theta\theta}^{(3)}, V_{\lambda\lambda}^{(3)}$ | $T_{rr}^{(3)}, T_{\theta\theta}^{(3)}, T_{\lambda\lambda}^{(3)}$ | $\eta^{(1)}$ |
| $V_{r\theta}^{(3)}, V_{r\lambda}^{(3)}, V_{\theta\lambda}^{(3)}$ | $T_{r\theta}^{(3)}, T_{r\lambda}^{(3)}, T_{\theta\lambda}^{(3)}$ | $\Theta^{(2)}$ |
| $V_{xx}^{(3)}, V_{yy}^{(3)}, V_{zz}^{(3)}$ | $T_{xx}^{(3)}, T_{yy}^{(3)}, T_{zz}^{(3)}$ | $(*)N^{(2)}$ |
| $(*)V_{xy}^{(4)}, (*)V_{xz}^{(4)}, (*)V_{yz}^{(4)}$ | $(*)T_{xy}^{(4)}, (*)T_{xz}^{(4)}, (*)T_{yz}^{(4)}$ | $\zeta_{Ell}^{(1)}$ |
| $W^{(1)}$ | $(**) \delta g^{(3)}$ | $(*) \zeta^{(2)}$ |
| $g^{(2)}$ | $\delta g_{sa}^{(1)}$ | |
| $g_{sa}^{(1)}$ | $\Delta g_{sa}^{(1)}$ | |
| $W_{rr}^{(1)}$ | $T_{rr}^{(1)}$ | |

Table 2: Statistical characteristics of differences between the reference values and those obtained using the “gradient approach in ellipsoidal coordinates”; used EGM2008 to degree 2190; gravity disturbances are in mGal ($1 \text{ mGal} = 10^{-5} \text{ m s}^{-2}$), second radial derivatives of disturbing potential in $1 \text{ E} = 10^{-9} \text{ s}^{-2}$ and CPU time in s.

| | δg_{sa} | | | | | T_{rr} | | | | |
|--------------------------------|-----------------|--------|--------|---------|----------|----------|---------|--------|---------|----------|
| | Min | Max | Mean | RMS | CPU time | Min | Max | Mean | RMS | CPU time |
| Reference – (0th order) | −101.661 | 46.375 | −6.712 | 14.6743 | 11 | −183.224 | 103.589 | −5.797 | 27.0377 | 11 |
| Reference – (0th + 1st order) | −23.622 | 38.123 | −0.285 | 3.3007 | 13 | −61.617 | 95.468 | −0.431 | 8.9100 | 14 |
| Reference – (0th to 2nd order) | −14.390 | 3.973 | −0.131 | 0.8870 | 16 | −41.712 | 14.055 | −0.290 | 2.7115 | 17 |
| Reference – (0th to 3rd order) | −1.749 | 4.564 | −0.016 | 0.2293 | 18 | −5.255 | 13.911 | −0.029 | 0.7280 | 18 |
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Table 3: Statistical characteristics of differences between the reference values and those obtained using the “gradient approach in spherical coordinates”; used EGM2008 to degree 2190; gravity disturbances are in mGal, second radial derivatives of disturbing potential in E and CPU time in s.

| | δg_{sa} | | | | | T_{rr} | | | | |
|---------------------------------|------------------------|-----------------------|------------------------|-----------------------|----------|------------------------|-----------------------|------------------------|-----------------------|----------|
| | Min | Max | Mean | RMS | CPU time | Min | Max | Mean | RMS | CPU time |
| Reference – (0th order) | −101.0 | 102.9 | −3.8 | 17.2 | 15 | −284.3 | 241.1 | −1.8 | 35.7 | 14 |
| Reference – (0th to 2nd order) | −11.0 | 9.6 | −2.4×10 ^{−2} | 1.4 | 19 | −30.1 | 26.8 | −2.1×10 ^{−2} | 4.3 | 18 |
| Reference – (0th to 5th order) | −0.1 | 0.1 | 7.8×10 ^{−5} | 1.8×10 ^{−2} | 27 | −0.4 | 0.4 | −1.8×10 ^{−4} | 6.0×10 ^{−2} | 25 |
| Reference – (0th to 10th order) | −1.3×10 ^{−5} | 9.8×10 ^{−6} | 3.8×10 ^{−9} | 1.4×10 ^{−6} | 38 | −4.2×10 ^{−5} | 3.2×10 ^{−5} | 1.0×10 ^{−10} | 4.6×10 ^{−6} | 38 |
| Reference – (0th to 15th order) | −1.5×10 ^{−10} | 1.2×10 ^{−10} | 2.4×10 ^{−14} | 1.5×10 ^{−11} | 49 | −4.7×10 ^{−10} | 3.6×10 ^{−10} | 1.2×10 ^{−13} | 4.7×10 ^{−11} | 46 |
| Reference – (0th to 20th order) | −4.8×10 ^{−11} | 4.3×10 ^{−11} | −1.8×10 ^{−14} | 4.8×10 ^{−12} | 60 | −9.8×10 ^{−11} | 1.2×10 ^{−10} | −1.3×10 ^{−14} | 9.4×10 ^{−12} | 56 |

Table 4: Computation time of the full disturbing tensor in the LNOF at 9 337 681 points using isGrafLab and the software presented by Eshagh and Abdollahzadeh (2012); used EGM2008 to degree 2190; K – the maximum order of the Taylor series; SFCM, ERA – the standard forward column method, the extended-range arithmetic, respectively (two of the three approaches for computing fnALFs in isGrafLab).

| | isGrafLab ($K = 3$) | isGrafLab ($K = 5$) | Eshagh and Abdollahzadeh |
|--------------|-----------------------|-----------------------|--------------------------|
| | SFCM / ERA | SFCM / ERA | |
| CPU time (h) | 0.7 / 1.0 | 1.1 / 1.5 | 13.4 |

Geopotential model and reference system selection

Global geopotential model of the Earth

Browse...

☒ Use maximum degree of GGM

GM of GGM (m3.s-2)

R of GGM (m)

nmin

nmax

Ellipsoid

3986004.415E+8

6378136.3

0

GRS80

Irregular surface selection

Type of the input coordinates:
☒ Ellipsoidal
☐ Spherical

Order of Taylor series:
3

Lat. min (°)

Lat. step (°)

Lat. max (°)

Irregular surface file

Browse...

Lon. min (°)

Lon. step (°)

Lon. max (°)

Structure of the input file:

☒ Matrix

☐ Column vector

Height above the reference surface (m)

Calculated parameters and output selection

Computation of fnALFs

☒ Export data

Display data settings

☒ Export report

Output folder and file

☐ Export data in *.mat

OK

Close

Figure 1: isGrafLab graphical user interface.

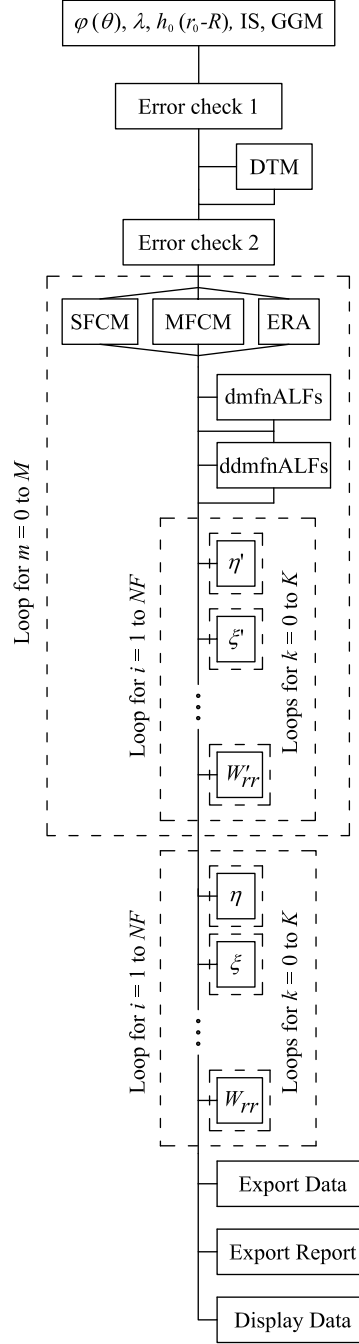


Figure 2: Flowchart of the isGrafLab. Explanation of the symbols and abbreviations in the flowchart: $\varphi(\theta), \lambda$ – ellipsoidal (spherical) coordinates of the evaluation points, $h_0(r_0 - R)$ – height of the regular surface above the reference ellipsoid (sphere), IS – input file containing the heights of the irregular surface, GGM – global geopotential model of the Earth; Error check 1 – the first error check of the input data; DTM – digital terrain model; Error check 2 – the second error check of the input data; SFCM, MFCM, ERA – computation of the modified fnALFs using the standard forward column method, the modified forward column method or the extended-range arithmetic approach, respectively; dmfnALFs – computation of the first-order derivatives of

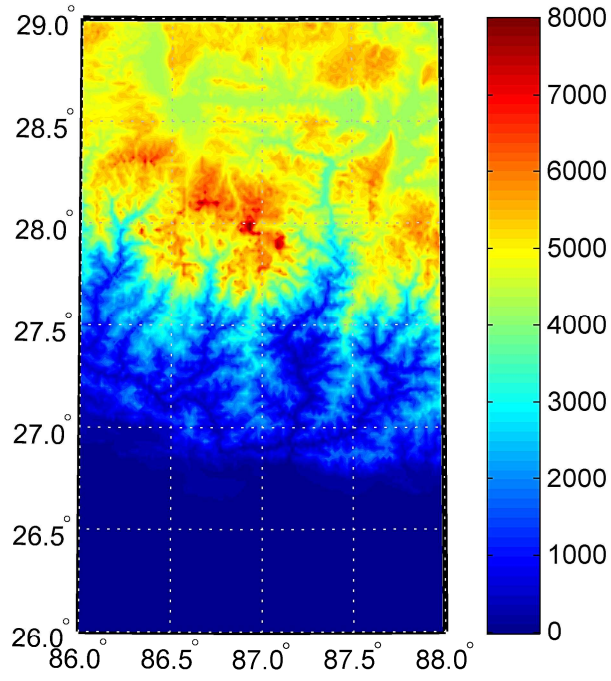


Figure 3: Topography of the Himalayas within the region $\varphi \in (26^\circ, 29^\circ)$, $\lambda \in (86^\circ, 88^\circ)$ at $1'$ spatial resolution. Units in metres.

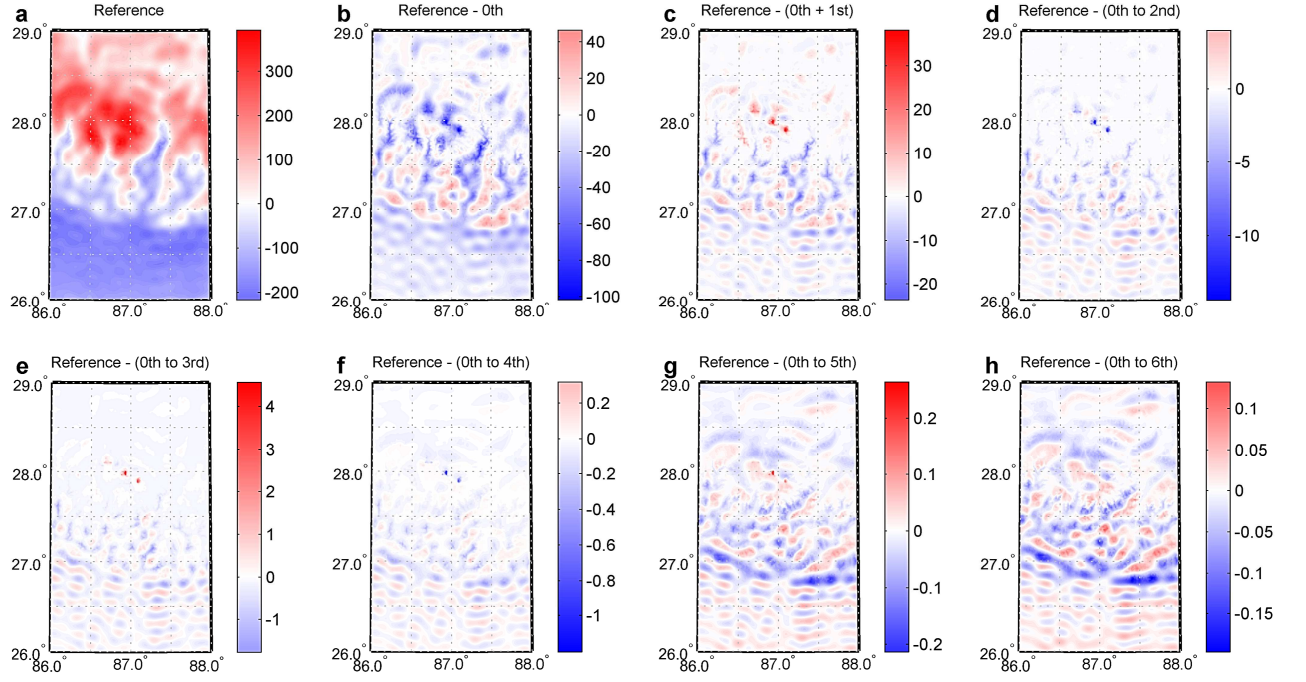


Figure 4: **a** Reference gravity disturbances in spherical approximation over the Himalayas, **b – h** differences between the reference and the approximated gravity disturbances in spherical approximation using the orders of the Taylor series $K = 0, \dots, 6$. Units in mGal.